



Frank Laboratory of
Neutron Physics

Possible UCN source at DNS-IV

Egor Lychagin



1970

1980

1990

2000

2010

2020



Frank Laboratory of Neutron Physics

Developing of experimental technic, of UCN sources, UCN detectors and etc.
Search reasons of abnormal large losses of UCN from traps

Discovering of hydrogen contamination of surfaces

τ_n 903±13

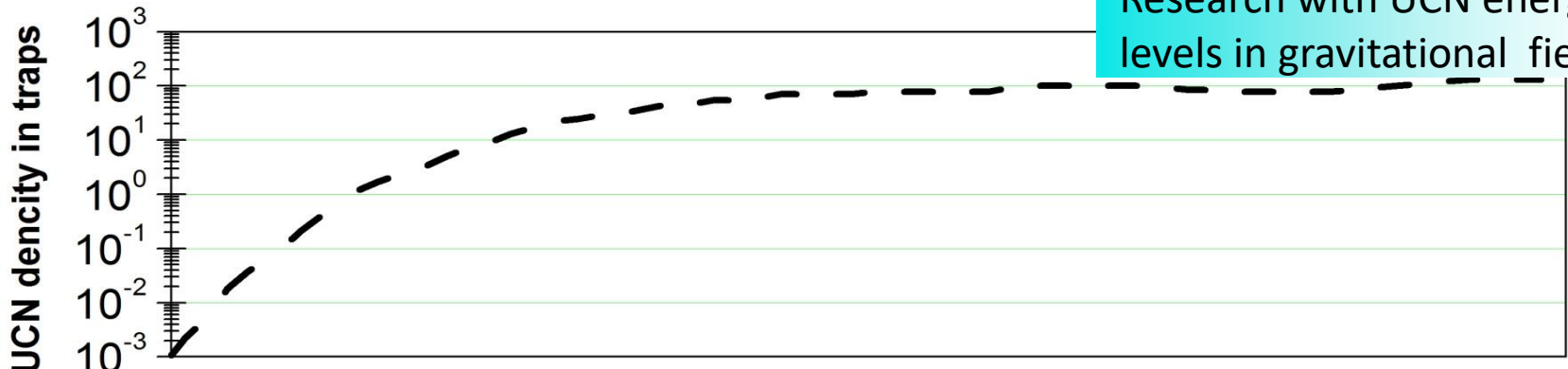
978±0.8

$|d_n| < 3 \cdot 10^{-24} \text{ e} \cdot \text{cm} \text{ (90\% C.L.)}$

$|d_n| < 9.7 \cdot 10^{-26} \text{ e} \cdot \text{cm} \text{ (90\% C.L.)}$

Discovering of UCN "small heating and cooling"

Research with UCN energy levels in gravitational field



To have breakthrough in the field it is need to develop and construct new powerful UCN sources.

Physical problems for UCN

- Neutron life time (electroweak interaction theory)
- Neutron EDM (T-invariance)
- Neutron wave properties (quantum mechanics, theory of relativity)
- Researches with gravitational levels
- Surface investigations
- Methodical developments

World UCN sources are in operation now

Neutron center	Neutron facility	UCN converter
ILL (PF2)	Steady state reactor, 58MW	VCN beam+Turbine
ILL (Sun-2)	Steady state reactor, 58MW	He-II
PSI	Spallation	SD ₂
UNIVERSITY OF MAINZ	TRIGA reactor, 250MW/25kW (pulse/average)	SD ₂
LANCE	Spallation	SD ₂
KEK	Spallation	He-II
NIST	Steady state reactor, 20MW	He-II

World UCN sources under construction

Neutron center	Neutron facility	UCN converter
North Carolina State University	Steady state reactor, 1MW	SD ₂
TUM	Steady state reactor, 20 MW	SD ₂
TRIUMF	Spallation	He-II

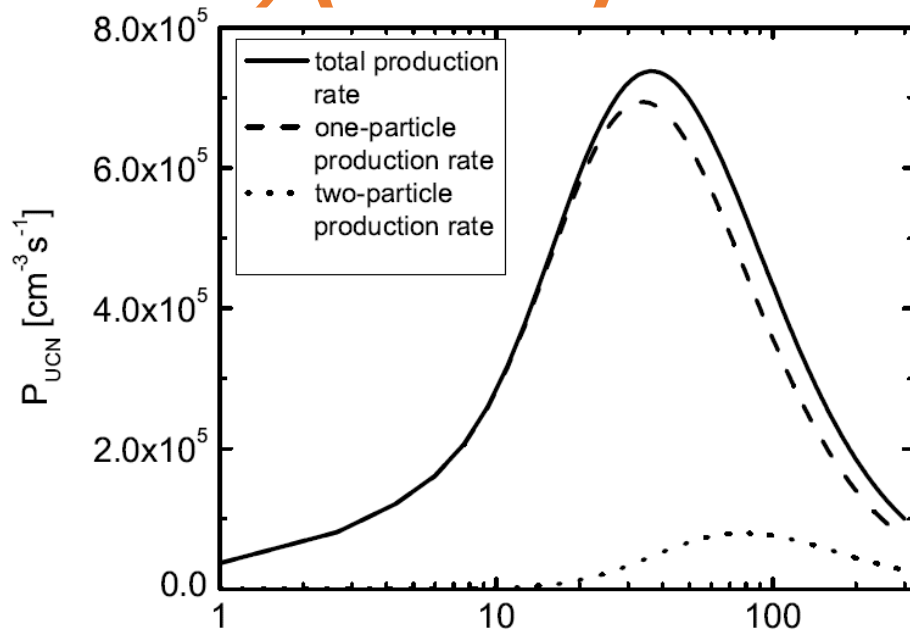
UCN sources projects under consideration

Neutron center	Neutron facility	UCN converter
SNS	Spallation	He-II
PNPI	Steady state reactor, 20 MW	SD ₂
PNPI	Steady state reactor, 100MW	He-II
PNPI	Steady state reactor, 100MW	He-II

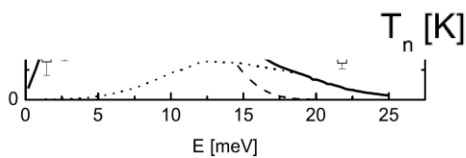
SD₂ (ortho)

vs

He-II



Temperature, K	<0.8
Free pass, cm	$\rightarrow \infty$
Energy, neV	105
Neutron spectrum structure T _n , K	6

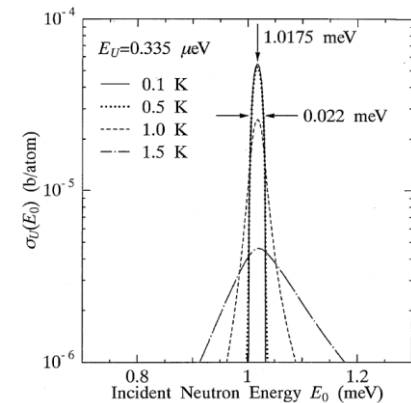


(0-150neV)

7×10⁵
(0-250neV)

action cross
tion

Production rate cm⁻³s⁻¹ for
neutron spectrum with
T=60K and F=1×10¹⁴ cm⁻²s⁻¹



(0-335neV)

4.5×10⁴
(0-250neV)

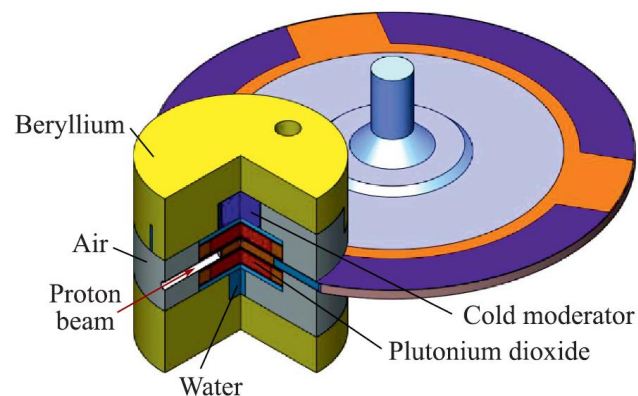
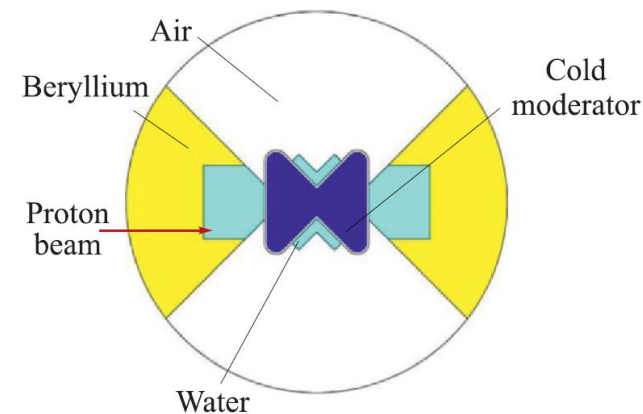
SD₂ (ortho) vs He-II

UCN life time in source is about 20ms.
Could have gain from pulse structure of
neutron flux at powerful bursts with
low repetition rate.

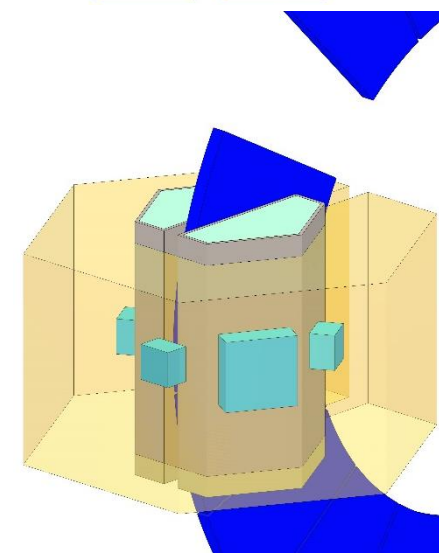
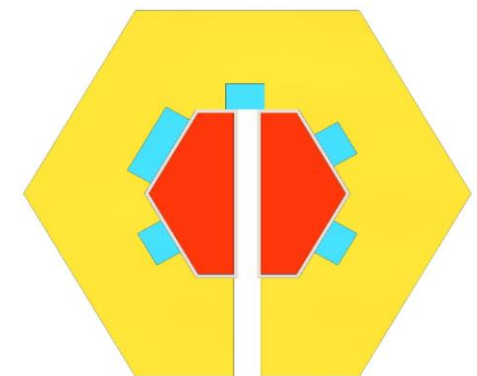
Under 0.8K allow accumulate UCN
density. It have no any gain from pulse
structure of neutron flux

At intense neutron source main
limitation in production rate origin from
cryogenicc problems.

DNS parameters

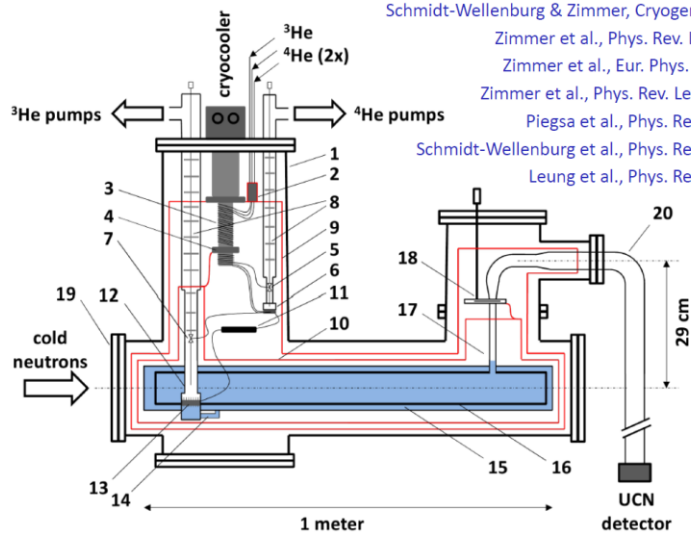
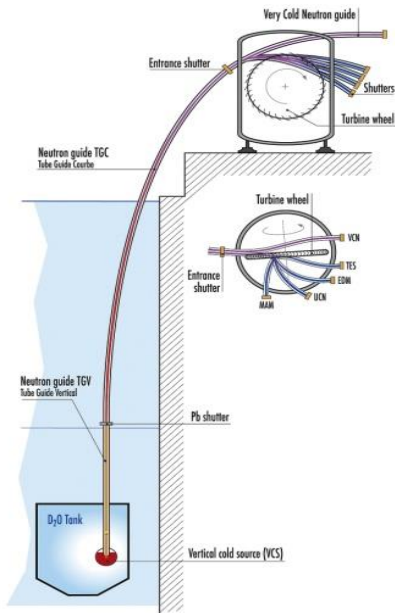


Average neutron flux density of thermal neutrons at the moderator surface, 1/(cm²s)	~1·10¹⁴
Peak neutron flux density of thermal neutrons at the moderator surface, 1/(cm²s)	~5·10¹⁶
Pulse width, ms	0.2
Repetition rate, Hz	10



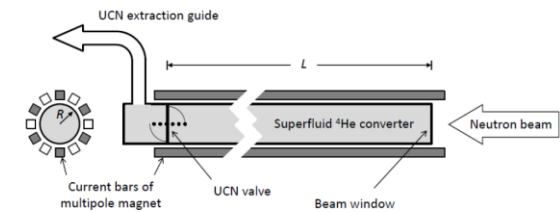
Sources at ILL

Source prototypes "SUN-1" & "SUN-2" (≥ 2004) window- and gap-less vertical UCN extraction



Schmidt-Wellenburg & Zimmer, *Cryogenics* **46** (2006) 799
 Zimmer et al., *Phys. Rev. Lett.* **99** (2007) 104801
 Zimmer et al., *Eur. Phys. J. C* **67** (2010) 589
 Zimmer et al., *Phys. Rev. Lett.* **107** (2011) 134801
 Piegsa et al., *Phys. Rev. C* **90** (2014) 015501
 Schmidt-Wellenburg et al., *Phys. Rev. C* **92** (2015) 024004
 Leung et al., *Phys. Rev. C* **93** (2016) 025501

SuperSUN



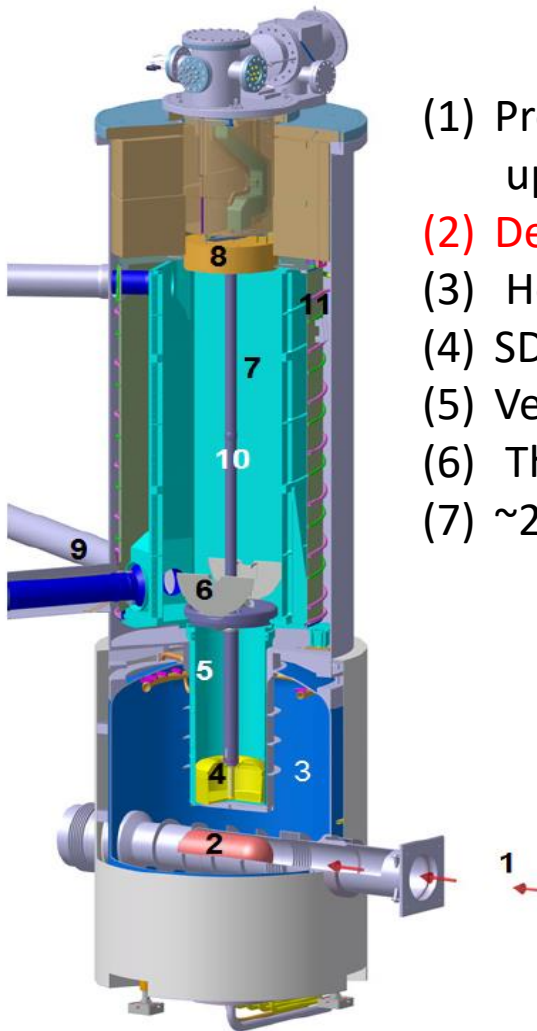
Advantages:

- in-situ UCN polarizer
- long storage lifetime \rightarrow high saturation UCN density
- weak dependence of ρ_{UCN} on wall quality

Converter volume: 12 litres
 UCN production rate: 10^9 s^{-1} ($E < 230 \text{ neV}$)
 UCN saturation number: 4×10^6 (2018, fomblin spectrum)
 2×10^7 (≥ 2019 , polarised, $E < 230 \text{ neV}$)

58 MW
 Heavy water tank
 Liquid D₂ cold source

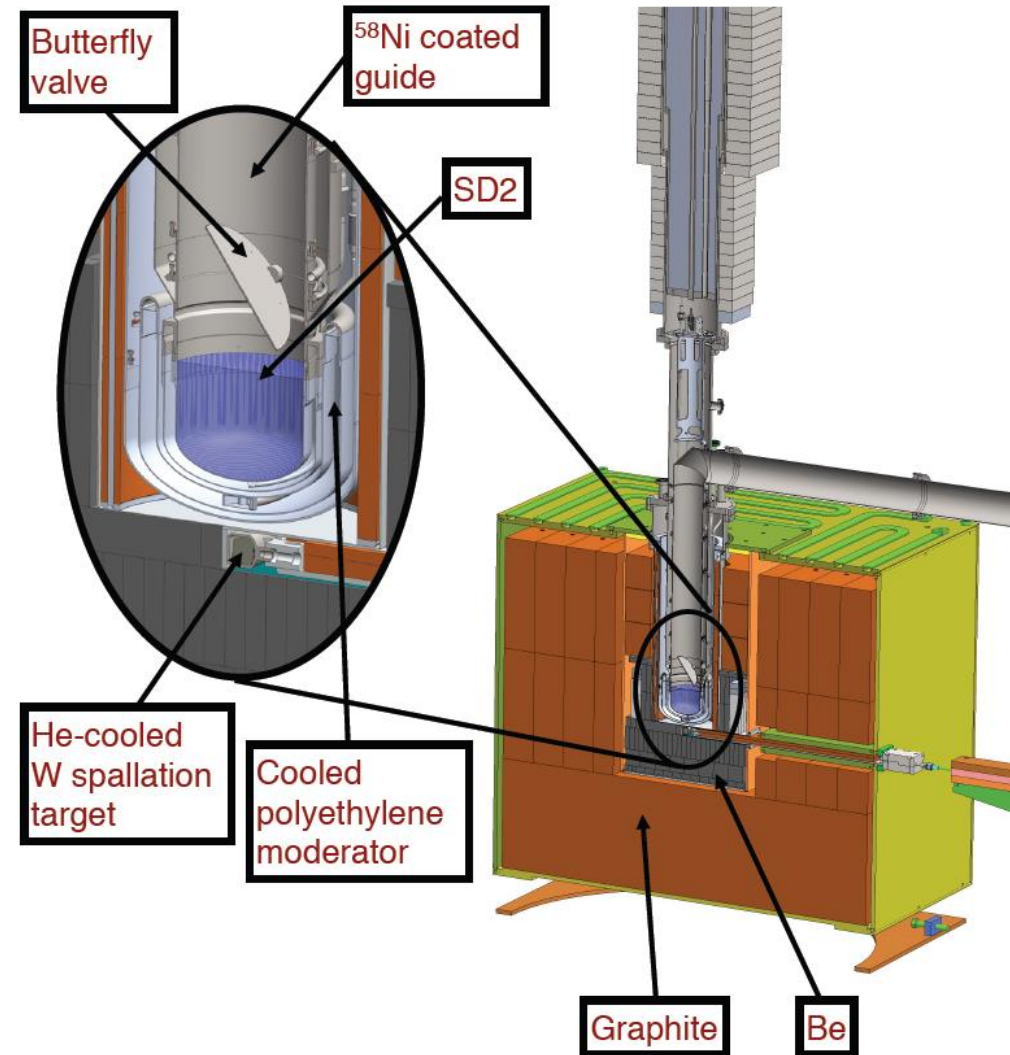
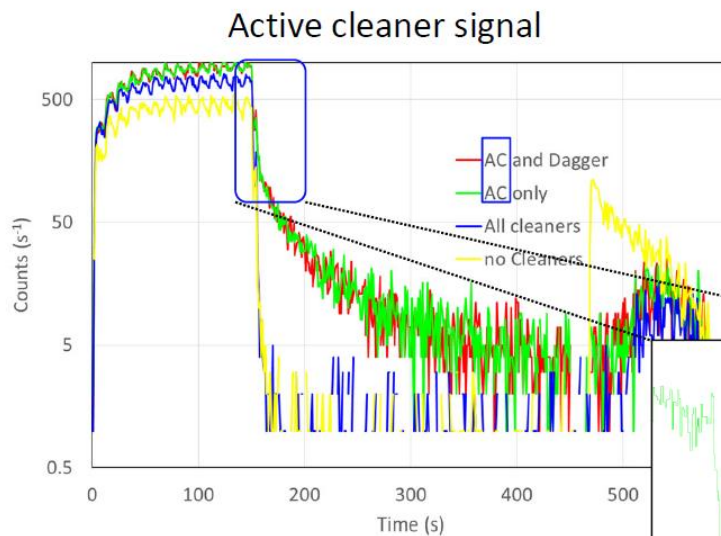
Source at PSI



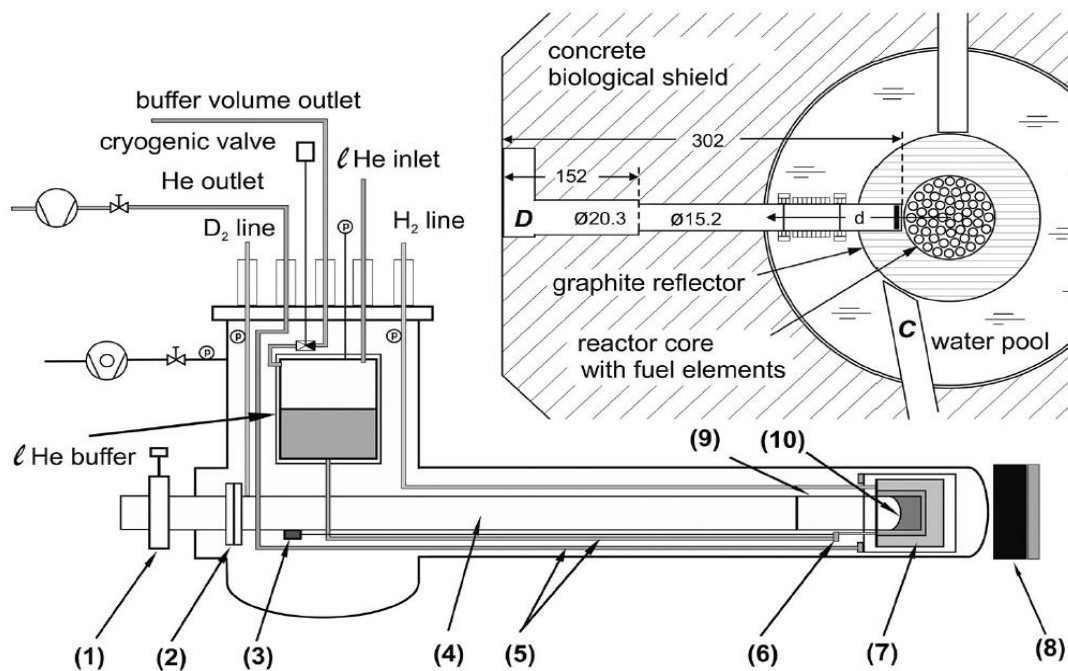
- (1) Proton beam 1.5×10^{16} protons/s (2.4mA) @590MeV
up to 8 s per 4-13.5 min
- (2) **Dedicated target**
- (3) Heavy water (D_2O) 3.5 m^3 moderator at a temperature of $\sim 300 \text{ K}$
- (4) SD_2 at 5 K (cold neutron flux of a few times $10^{13} \text{ n/cm}^2/\text{s}$ inside)
- (5) Vertical guide
- (6) The big shutter
- (7) $\sim 2 \text{ m}^3$ large UCN storage vessel.

Source at LANCE

- Proton beam $45\mu\text{C}$ during 0.45s (in 10 pulses with $625\mu\text{s}$ width and 20 Hz repetition) with gap 5 s
- **Dedicated target**
- SD_2 at 5 K ($\sim 1800\text{ cm}^3$)
- Vertical guide
- The big shutter



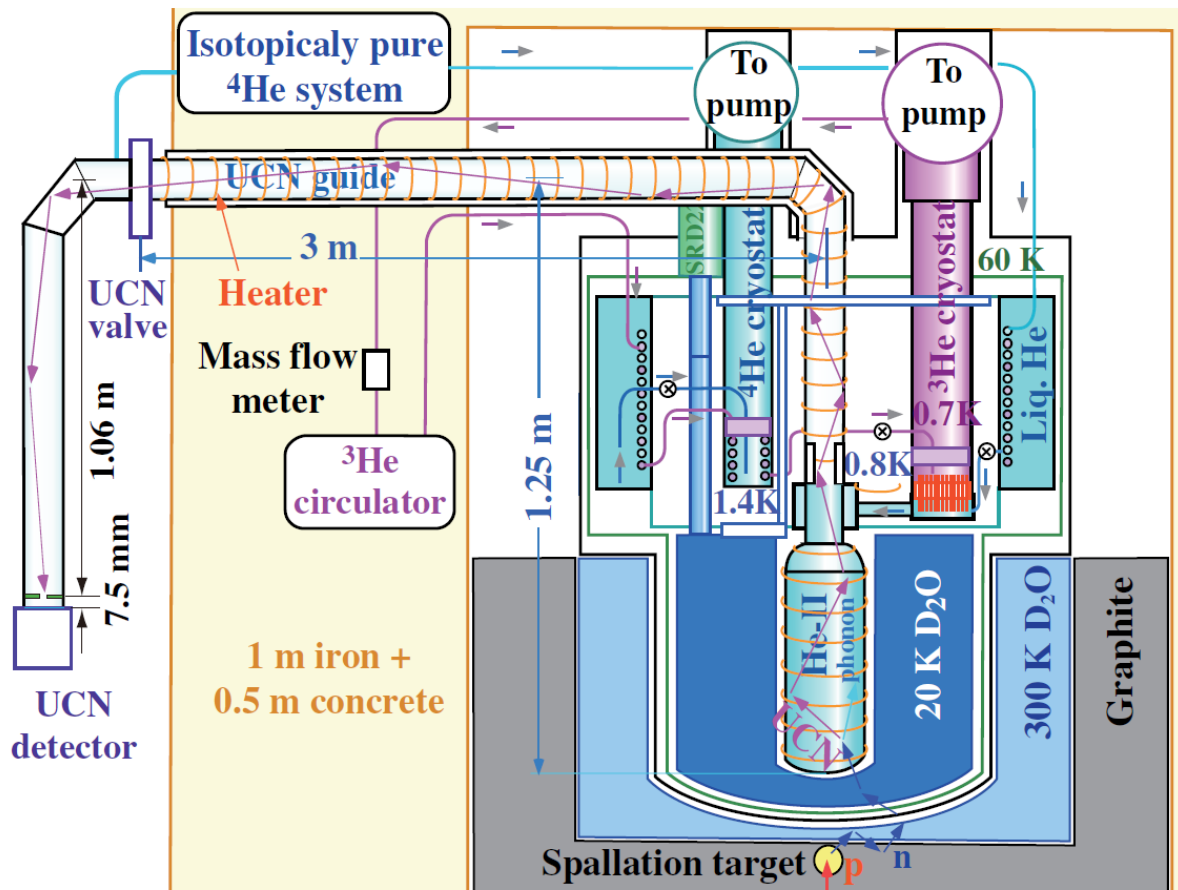
Source at University of Mainz



- TRIGA reactor 250 MWs (in pulse)
- Pulse 30ms per 5 min
- Solid methane/para H₂ pre-moderator
- SD₂ at 6 K (~160 cm³)

(7) premoderator (H₂, D₂ or CH₄), (8) graphite/bismuth stopper, (10) nose with SD₂ converter.
 Inset: Scale drawing (measures in cm) of the horizontal section at reactor TRIGA Mainz with focus on the radial beamport D

Source at KEK



- Proton beam 400 MeV, 1 μ A during 100 s
- Dedicated target
- Steady state regime

It is stationary neutron source indeed

Development at North Carolina State University

PULSTAR Reactor 1MW

- Heavy water tank (300K)
- Solid methane pre-moderator (40K)
- SD_2 at 5 K

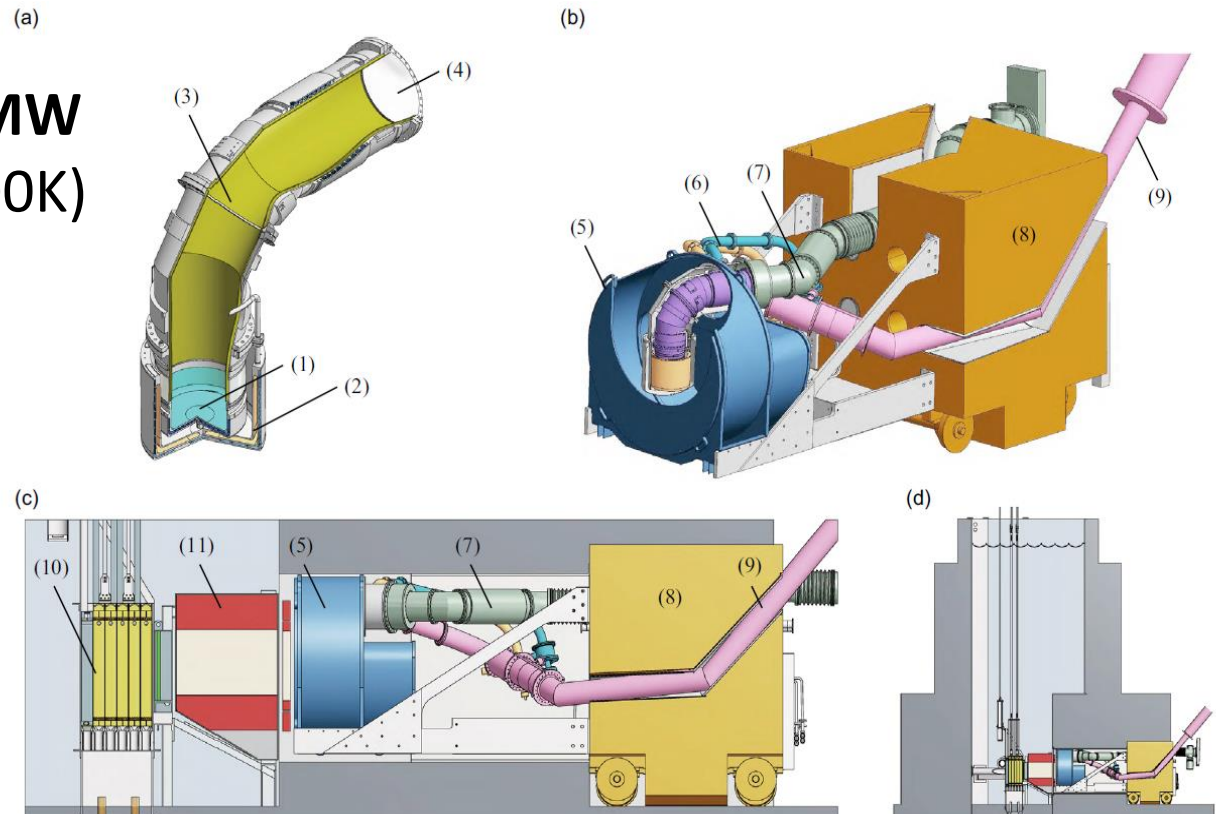
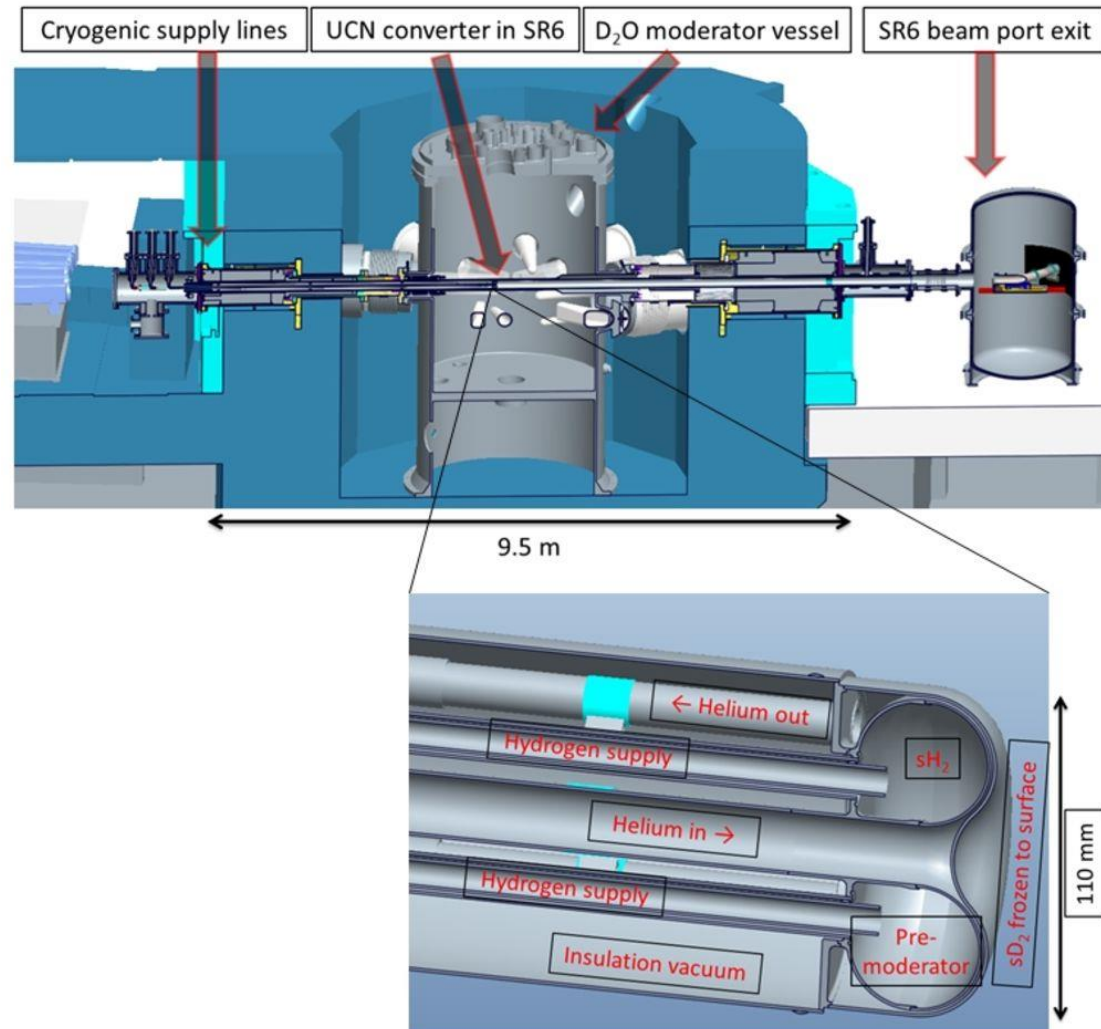


Fig. 1. A cutaway engineering drawing of the UCN source cryostat, as envisioned installed in the reactor thermal column, is shown in (a), including the (1) deuterium container, (2) methane container, (3) UCN guide, and (4) UCN window foil. The UCN source assembly mounted on the shielding door is shown in (b), including the (5) heavy-water tank, (6) deuterium and methane gas inlets, (7) UCN guide, (8) thermal column shield door, and (9) liquid helium transfer line. The heavy-water tank and the cryostat vacuum jacket have been cut away to show the cryostat. A cross-section of the reactor showing the source assembly in the thermal column is shown in (c) and (d), where (10) is the reactor core and (11) is the neutron transport system. In all drawings, neutrons from the reactor core enter from the left and UCN exit to the right of the shielding door.

Development at TUM

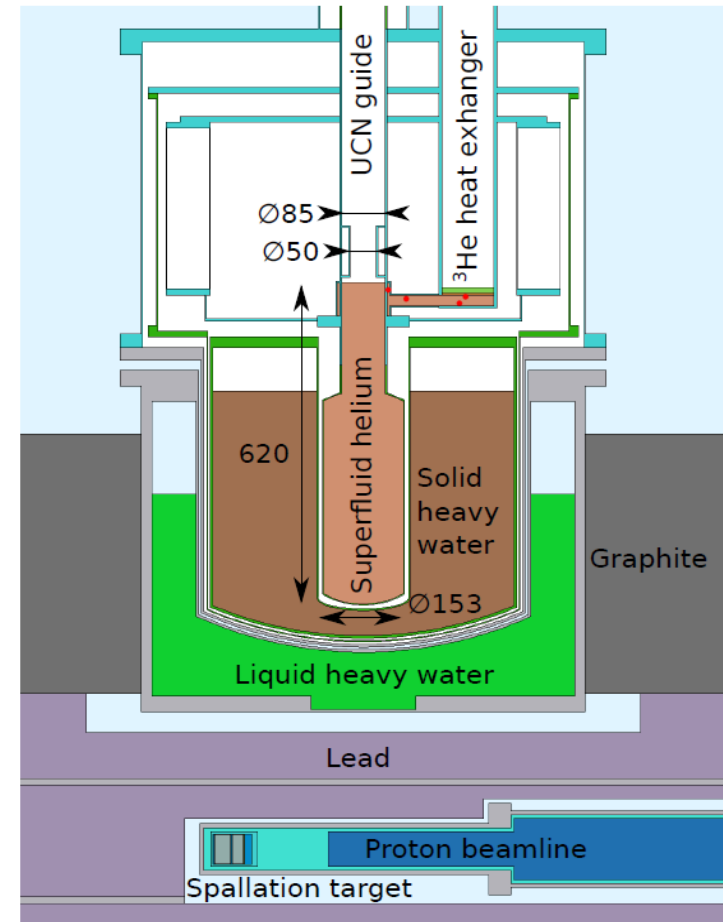
FRM II reactor 20 MW

- sD_2 at 5 K ($\sim 250 \text{ cm}^3$)
- Pre-moderator volume (sH_2) 250 cm^3
- Converter volume (sD_2) 250 cm^3

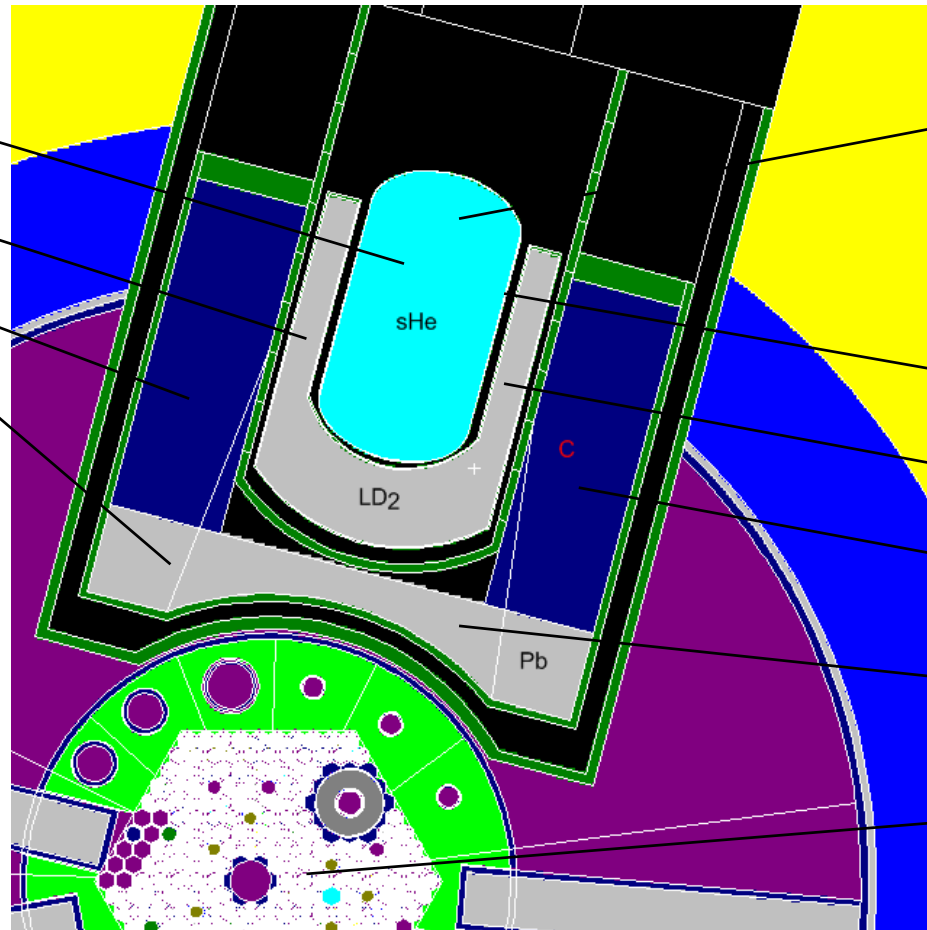


Development at TRIUMF

- Spallation at tungsten by protons 482 MeV, 40 μA
- Pulsed 1ms, 33 Hz (“quasi-stationary”)
- Dedicated target



Project for VVR-M at PNPI



$\rho_{UCN} = 10^4 \text{ cm}^{-3}$ ($\tau = 10 \text{ c}$)
 $\Phi = 4.5 \cdot 10^{12} \text{ n}/(\text{cm}^2 \text{c})$
 $\Phi(\lambda = 9 \text{ \AA}) = 3 \cdot 10^{10} \text{ n}/(\text{cm}^2 \text{c})$
 $Q_{\text{He}} = 6 \text{ W}$

19 W

Al, $Q_{\text{Al}} = 13 \text{ W}$

LD₂, $Q_{\text{LD}_2 + \text{Al}} = 100 \text{ W}$

C, $Q_{\text{C}} = 700 \text{ W}$

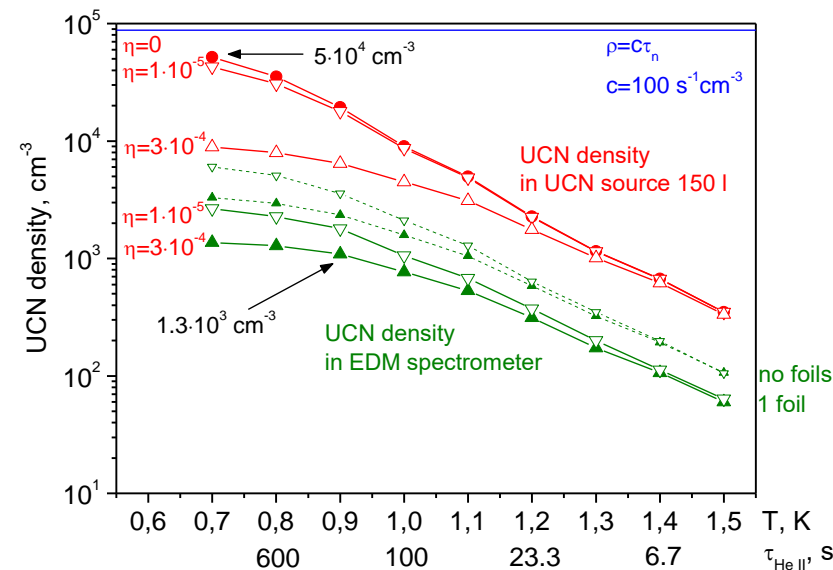
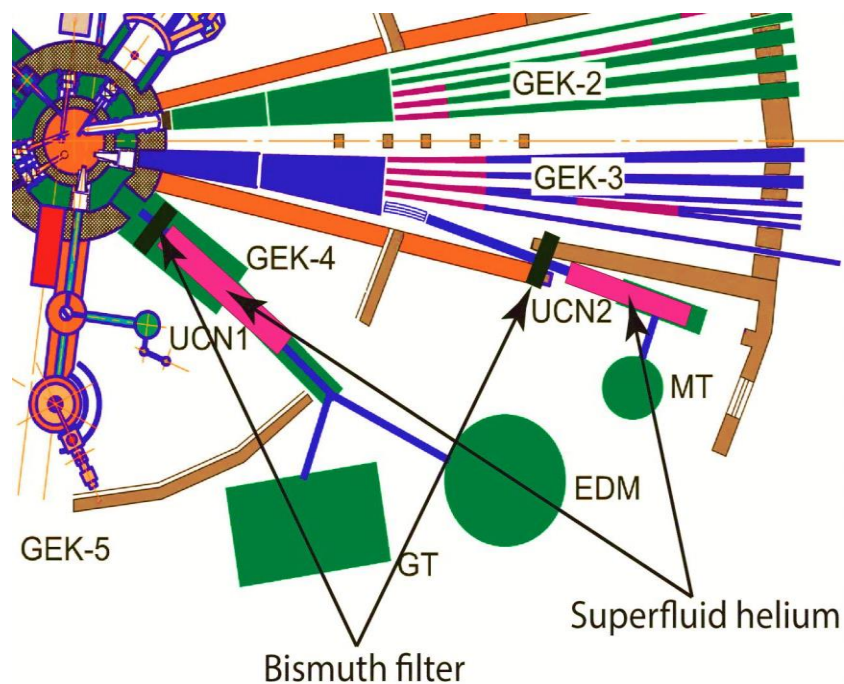
Pb, $Q_{\text{Pb}} = 15 \text{ kW}$

$\Phi = 10^{14} \text{ n}/(\text{cm}^2 \text{c})$

$Q = 15 \text{ MW}$

Project for PIK at PNPI

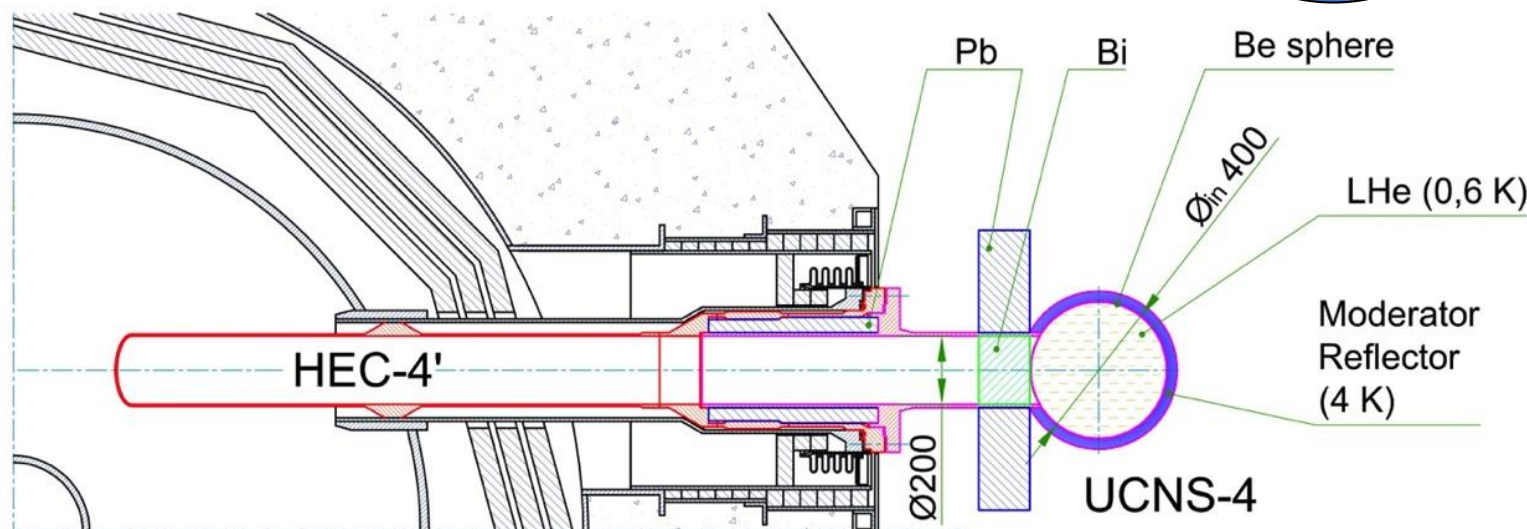
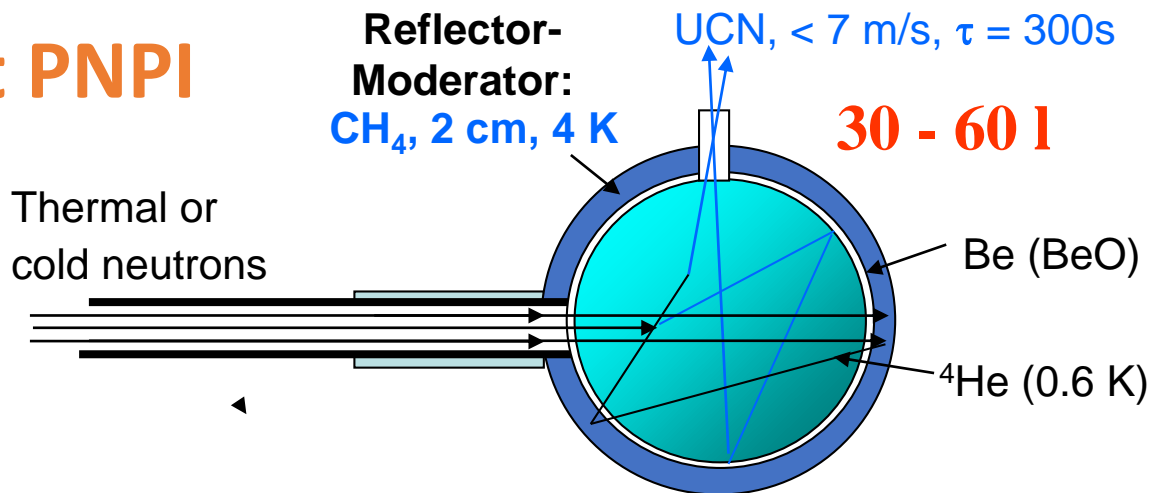
The general scheme of a complex of experimental installations for carrying out research of fundamental interactions with UCN at PIK reactor



**Extracted beam
Need cold source**

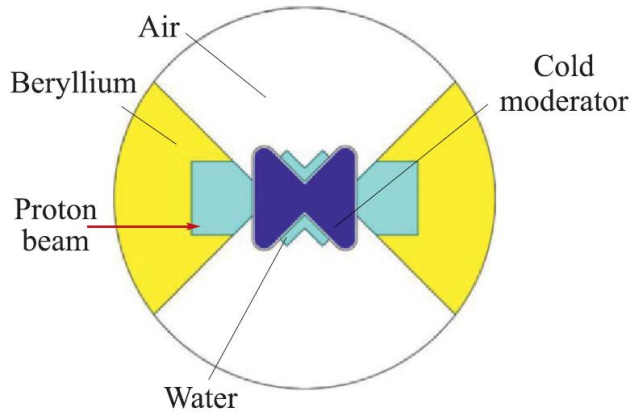
The thermal load will be much higher at DNS due to hydrogen moderators. The gamma flux should be decreased, or heavy water moderator should be installed (neutron flux will be suppressed)

Project for PIK at PNPI

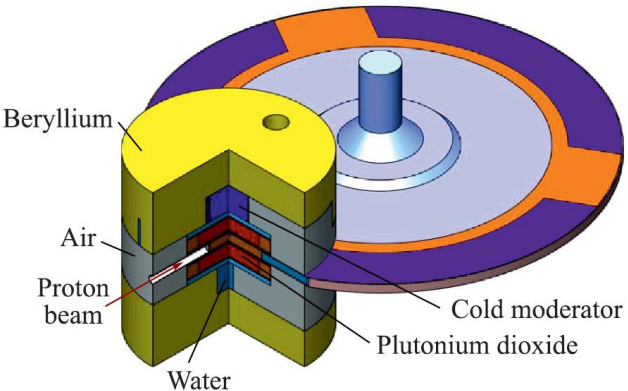
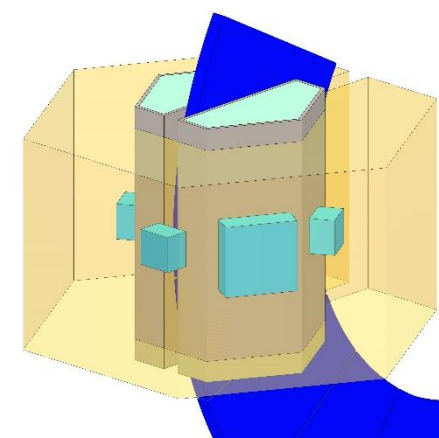
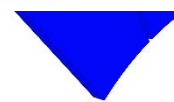
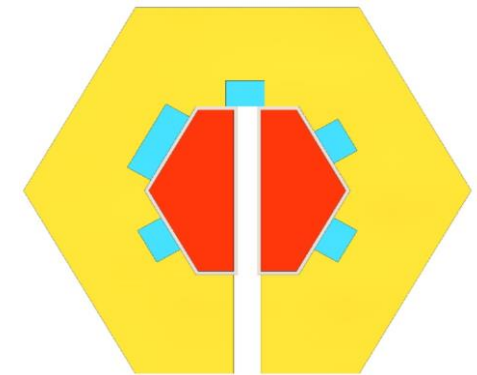


UCN density $\sim 10^5 \text{ cm}^{-3}$ UCN flux $\sim 10^7 \text{ c}^{-1}$

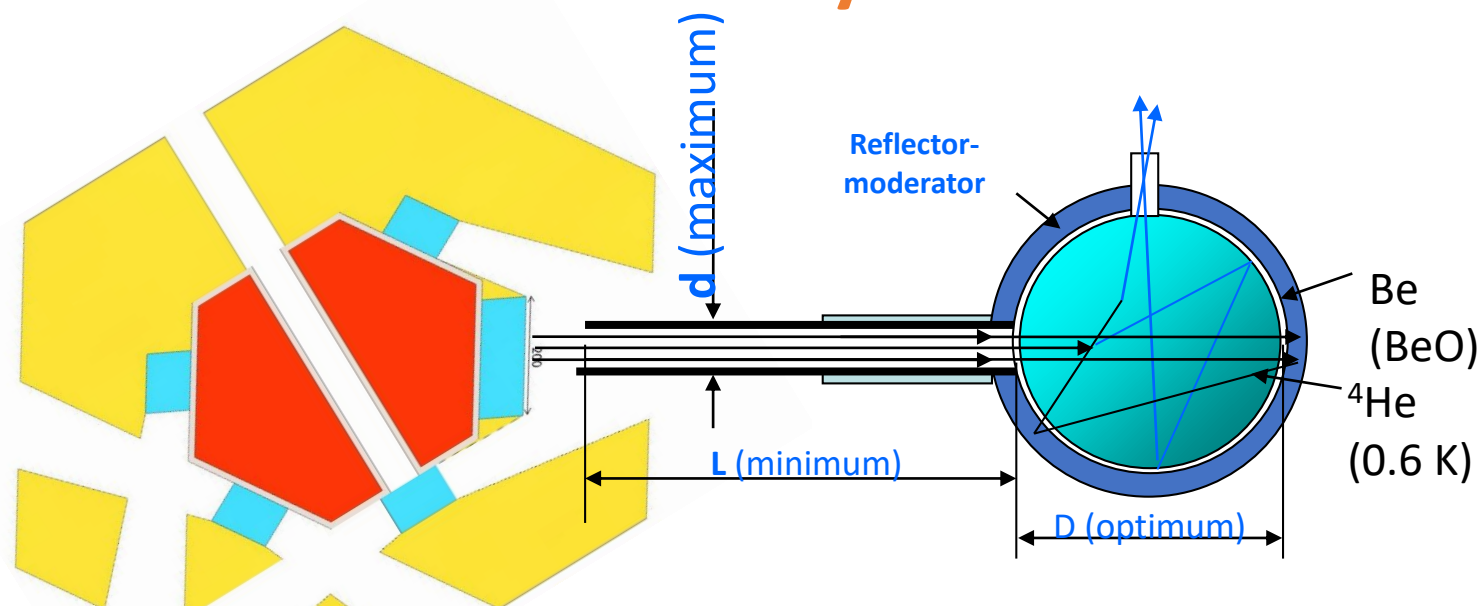
DNS parameters



Average neutron flux density of thermal neutrons at the moderator surface, 1/(cm²s)	~1·10¹⁴
Peak neutron flux density of thermal neutrons at the moderator surface, 1/(cm²s)	~5·10¹⁶
Pulse width, ms	0.2
Repetition rate, Hz	10



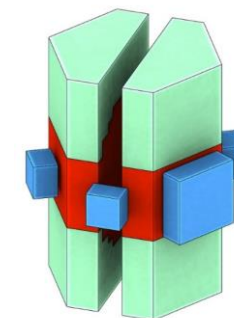
Possibility for DNS



- Will need thick ($\sim 250\text{mm}$) Bi gamma filter
- It is better to have a cold source (less flux suppression by the filter)
- Rough estimation: **production rate** $\cdot 3 \times 10^5 \text{ n/s}$ (less than at PF2 ILL); **density** $\sim 3 \times 10^3 \text{ n/cm}^3$ (100 times more than at PF2 ILL) with $5 \times 10^{13} \text{ n/cm}^2/\text{s}$ at the moderator surface (20 times less than at the PIK)

Calculation of neutron and gamma fluxes needed to define the source parameters
Some questions (trap material, effective extraction) are open.

Possibility for DNS



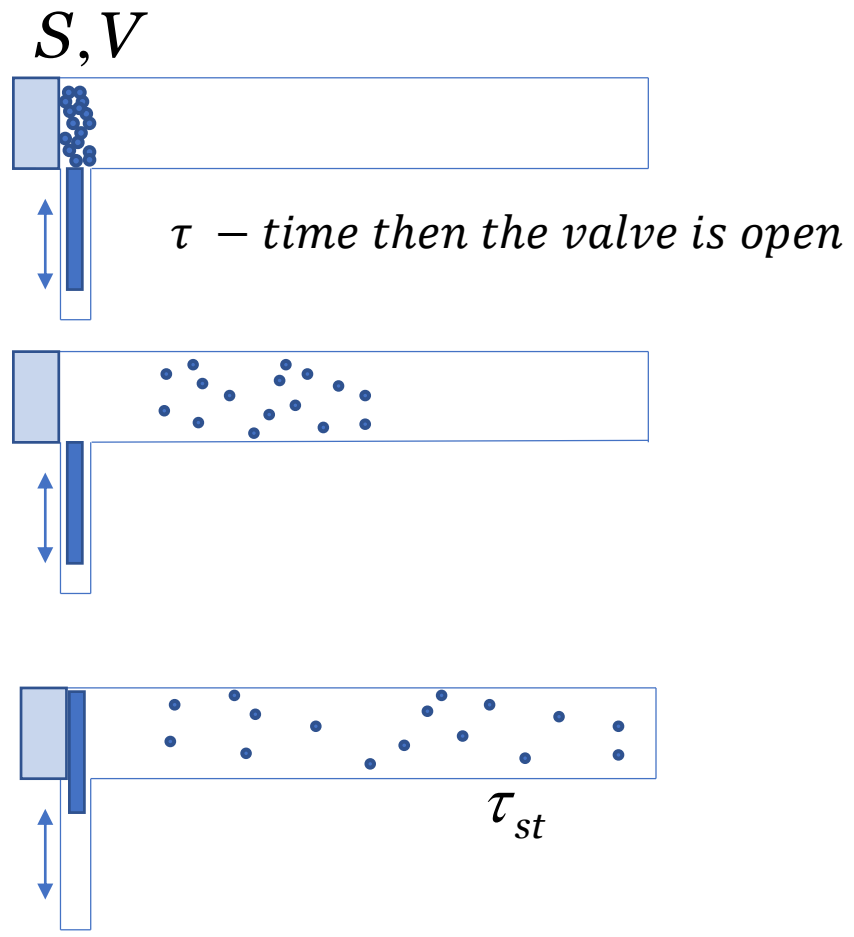
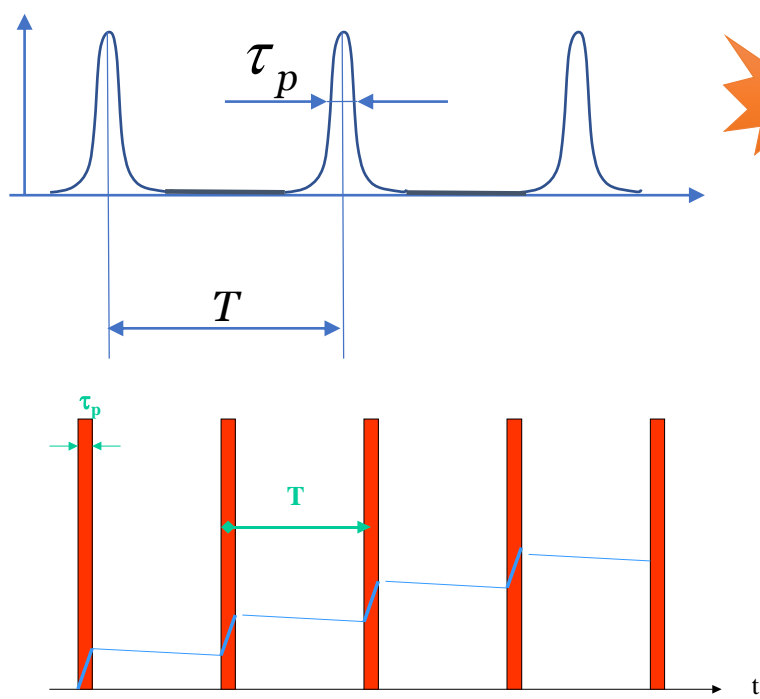
Cold moderator + SD_2

With $1 \times 10^{13} \text{ n/cm}^2/\text{s}$ of cold neutron spectra in 500 cm^3 of SD_2 the production rate will be about $\cdot 1 \times 10^8 \text{ n/s}$ (100 times more than at ILL).

- In addition it is a source of VCN
- The VCN source brightness could be increased by VCN reflector but TOF will be lost (due to the VCN pulse widening)

Calculation of neutron flux and heat relies needed to define the source parameters

Pulse density accumulation?



Pulse density accumulation?

At equilibrium case:

$$\frac{\rho_{imp} v}{4} S \tau_p = \frac{\rho v}{4} S \tau + \rho V (1 - e^{-\frac{T}{\tau_{st}}})$$

$$\rho = \frac{\rho_{imp}}{\frac{\tau}{\tau_p} + \frac{4V}{Su} \frac{T}{\tau_{st} \tau_p}} \quad \frac{4V}{Su} = \tau_{emp}$$

$$\rho = \frac{\rho_{imp}}{\frac{\tau}{\tau_p} + \frac{\tau_{emp} T}{\tau_{st} \tau_p}}$$

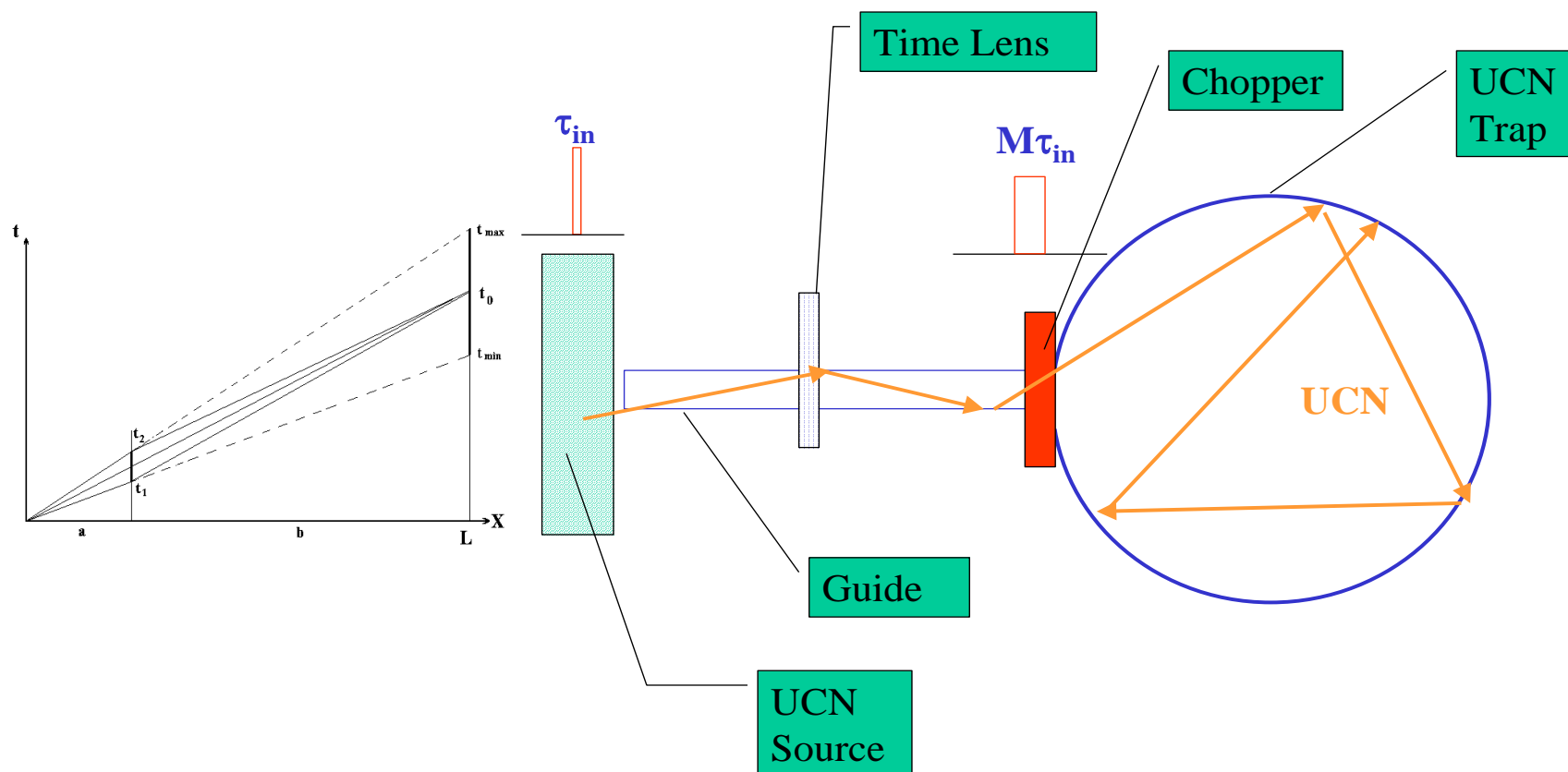
$$\rho = \begin{cases} \rho_{imp} & \text{if } \frac{\tau_{st}}{\tau_{emp}} \ll \frac{T}{\tau_p} \\ \rho_{imp} \frac{\tau_p}{T} & \text{if } \frac{\tau_{st}}{\tau_{emp}} \sim 1 \end{cases}$$

Gain factor for DNS is 500
But technical realization is
questionable.

(For ESS gain factor is 30)

Pulse density accumulation?

A.I.Frank and R.Gähler. Phys. of Atomic Nuclei, 63, (2000) 545



Conclusions:

1. Solid deuterium source looks reasonable for DNS.
2. The extensive calculations (fluxes, heat release etc.) are needed to define the source parameters.
3. Need relatively wide neutron channel without the shatter or with special shatter at any case.
4. It is good to have relatively thin biological shielding (vertical extraction of UCN possible).
5. Accumulation of pulsed density looks attractive but difficult from technical point. The time focusing of UCN "cloud" in mirror neutron guide need search and investigation of «time lens».
6. We should start at IBR-2.

Thank you for attention